

Testing gravity on accelerators

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- 1506.08063 (PLB)** Gravitational mass of relativistic matter and antimatter
- 1506.01963 (PLB)** Testing general relativity on accelerators
- 1604.04486 (PRL)** Comment on “Testing Planck-Scale Gravity with Accelerators”
- 1508.04377 (Sci.Rep.)** Gravitational mass of positron from LEP synchrotron losses

Disclaimer: This work was done as a private venture and not in the author's capacity as an employee of the Jet Propulsion Laboratory, California Institute of Technology.

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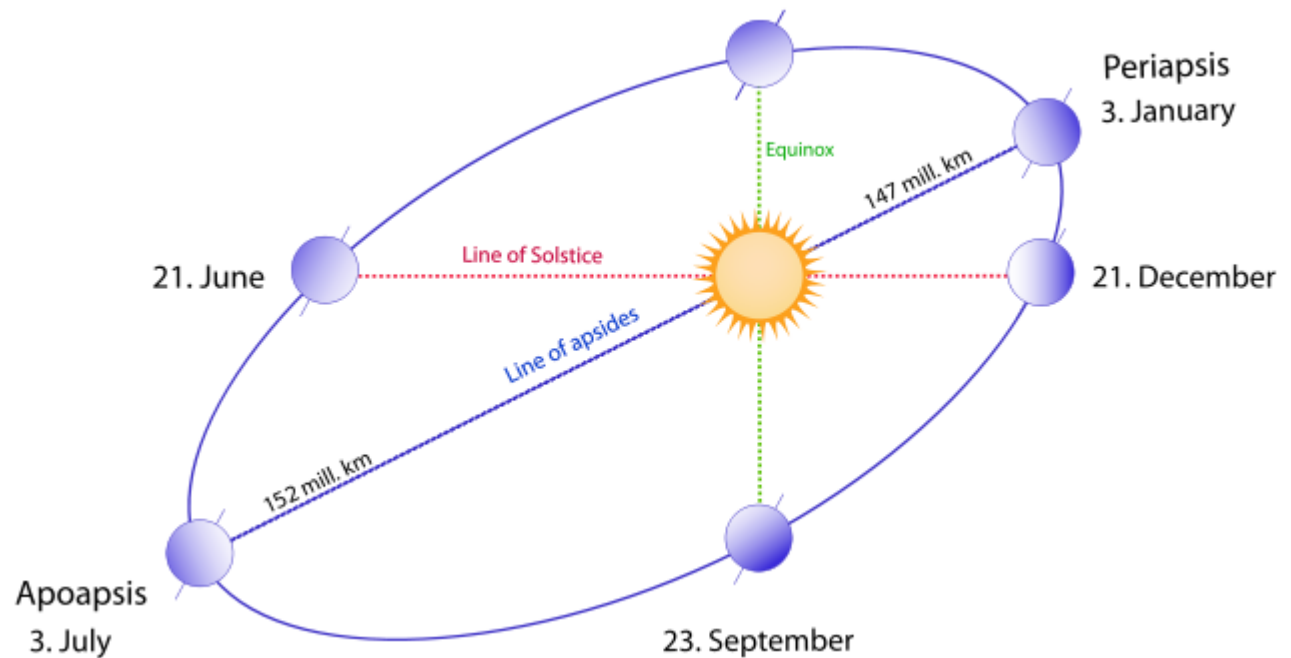
Motivation

Tests of gravity at high energies

Antimatter gravity $\left(-65 < m_g^{\bar{H}} / m^{\bar{H}} < 110\right)$

How?

**Perform tests
on the isotropic
Lorentz
violation on two
different days
of the year.**



Theory in brief

Gravitational field around the accelerator:

$$ds^2 = \mathcal{H}^2 dt^2 - \mathcal{H}^{-2}(dx^2 + dy^2 + dz^2) \quad \text{where} \quad \mathcal{H}^2 = 1 + 2\Phi$$

For a massive particle (in our case ultrarelativistic electron or positron)

$$\Phi_m = \Phi \frac{m_{e,g}}{m_e}, \quad \mathcal{H}_m^2 \equiv 1 + 2\Phi_m$$

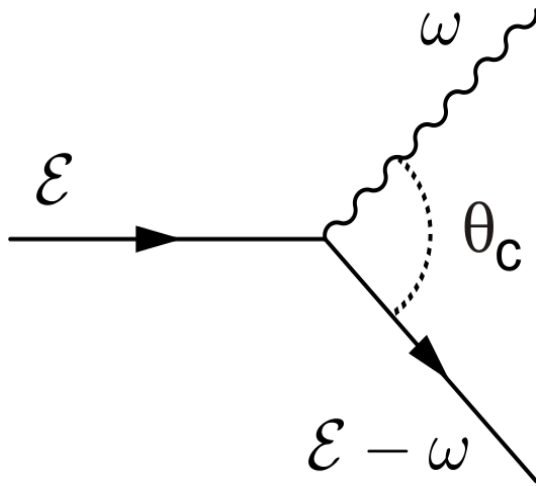
which will modify the dispersion relation of the particle and the relation between energy and mass (we assume the speed of light to be universal)

$$\mathbf{p}^2 = (1 + 2\kappa) (\mathcal{E}^2 - m_e^2), \quad \mathcal{E} = \frac{m_e \mathcal{H}^{-1} \mathcal{H}_m}{\sqrt{1 - \mathcal{H}^4 \mathcal{H}_m^{-4} \mathbf{v}^2}},$$

where $\kappa = 2\Phi \Delta m_e / m_e$, $\Delta m_e = m_{e,g} - m_e$.

Imagine that $|\kappa| < \kappa_{1,2}$ **for two experiments, then** $\left| \frac{\Delta m_e}{m_e} \right| < \frac{\kappa_1 + \kappa_2}{2\Delta\Phi}$

1. Vacuum Cherenkov radiation



Threshold energy:
$$\mathcal{E}_{\text{th}} = \frac{m_e}{\sqrt{-2\kappa}}$$

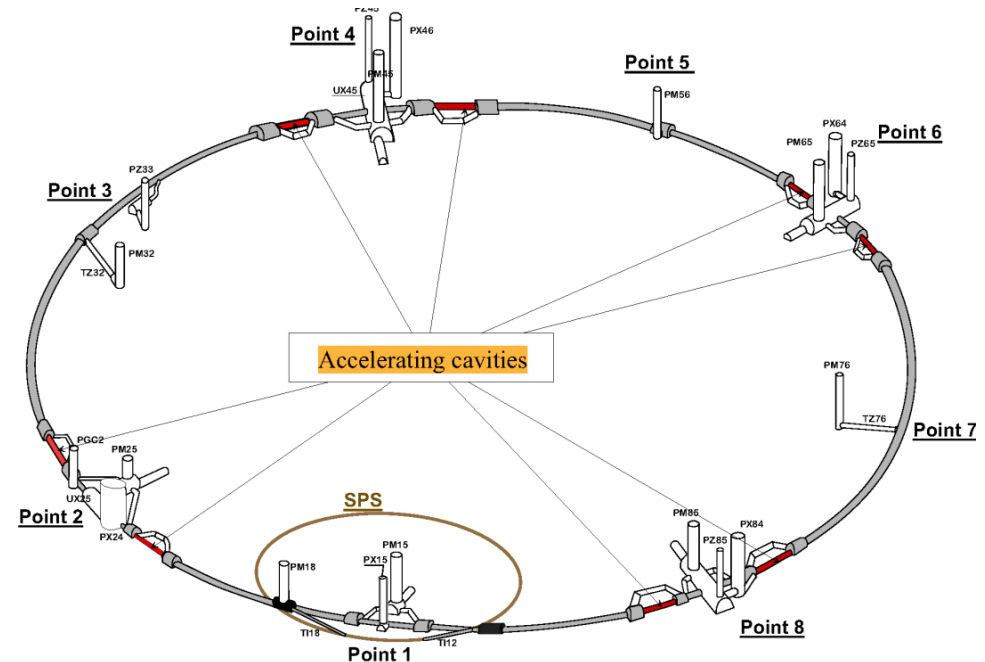
Emission rate:
$$\Gamma_C = \alpha m_e^2 \frac{(\mathcal{E} - \mathcal{E}_{\text{th}})^2}{2\mathcal{E}^3}$$

Let us take $E = 104.5$ GeV electrons and positrons at LEP.

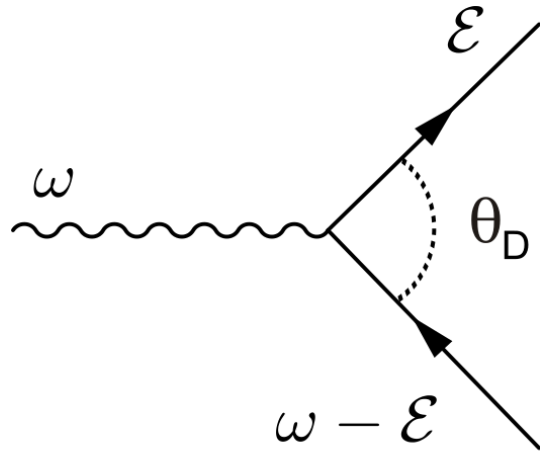
$$\mathcal{E}_{\text{th}} = 100 \text{ GeV}$$

Compare: 1.2cm (decceleration distance) vs 6 km (approximate distance between accelerating RF systems).

$$\kappa > -1.3 \times 10^{-11}$$



2. Photon decay



Threshold energy:

$$\omega_{\text{th}} = \sqrt{\frac{2}{\kappa}} m_e$$

Decay rate:
$$\Gamma_D = \frac{2}{3} \alpha \omega \frac{m_e^2}{\omega_{\text{th}}^2} \left(2 + \frac{\omega_{\text{th}}^2}{\omega^2} \right) \sqrt{1 - \frac{\omega_{\text{th}}^2}{\omega^2}}$$

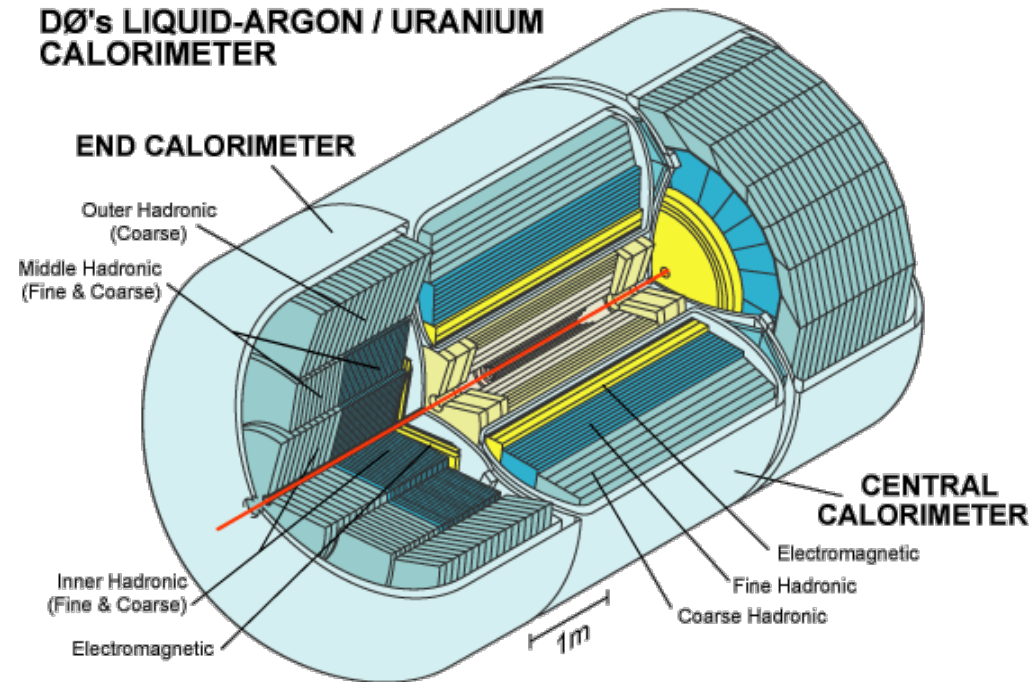
Let us take $E = 340.5 \text{ GeV}$ photons at Fermilab's Tevatron.

$$\omega_{\text{th}} = 300 \text{ GeV}$$

Compare: 0.1 mm (decay distance)
vs 78 cm (minimal path from interaction
point to the central calorimeter of D0
detector).

$$\kappa < 5.8 \times 10^{-12}$$

DØ's LIQUID-ARGON / URANIUM
CALORIMETER



3. Synchrotron radiation

Radiation power without gravity

$$P = \frac{2}{3} \frac{e^2 \dot{\mathbf{v}}^2}{c^3} \left(\frac{\mathcal{E}}{m_e} \right)^4$$

Modification of the gamma-factor leads to

$$\Delta P/P = 4\kappa\gamma^2$$

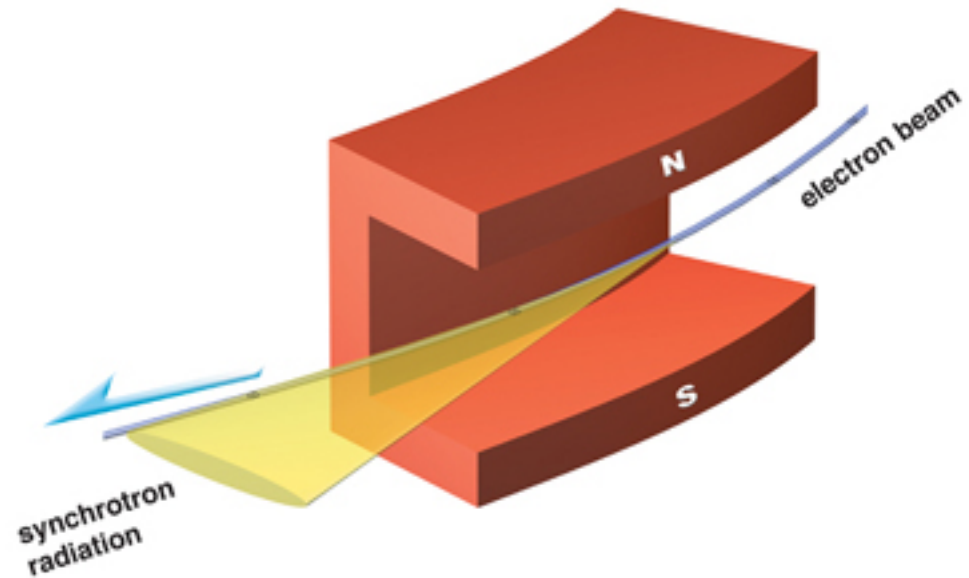
LEP E = 80 GeV electrons and positrons.

Energy was estimated by 3 methods: NMR and flux-loop magnetic field measurement; spectrometry; synchrotron tune vs RF voltage fit.

$$Q_s^4 = \left(\frac{\alpha_c h}{2\pi} \right)^2 \left\{ \frac{g^2 e^2 V_{RF}^2}{E^2} + M g^4 V_{RF}^4 - \frac{U^2}{E^2} \right\}$$

One can reinterpret it as a fit to U and possible uncertainty in the synchrotron losses

$$|\kappa| < 9 \times 10^{-15} \quad \text{for two experiments (13 Aug \& 15 Sep 1999)}$$



Taking the two-sided bound $\kappa^- < \kappa < \kappa^+$ for two potentials, Φ and $\Phi + \Delta\Phi$ (e.g., both the vacuum Cherenkov radiation and photon decay were absent during the experiment), one can easily derive

$$\kappa^- - \kappa^+ < -2\Phi_{\odot} \frac{\Delta d_{SE}}{d_{SE}} \frac{\Delta m_e}{m_e} < \kappa^+ - \kappa^- ,$$

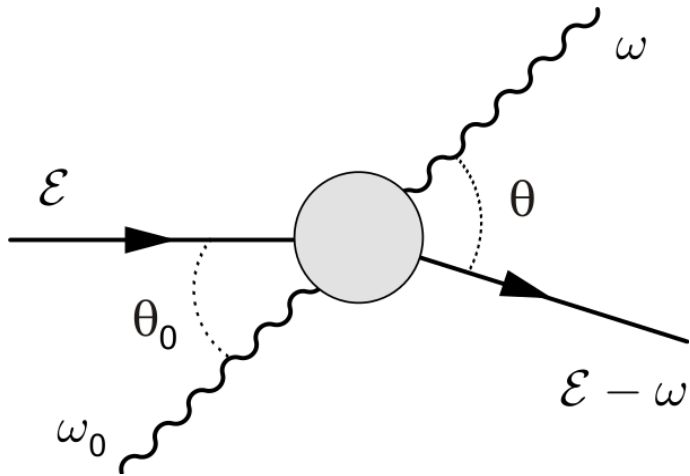
where Δd_{SE} is the variation of the distance between Sun and Earth, d_{SE} , due to the eccentricity of the Earth's orbit, and obtain $|\Delta m_e/m_e| < 0.0389$, i.e., a 4% limit on the possible deviation.

One can also relate the fractional deviation in the masses and the fractional uncertainty in the measured synchrotron radiation power in two experiments,

$$\left| \frac{\Delta m_e}{m_e} \right| < \frac{|\Delta P/P|_1 + |\Delta P/P|_2}{8\gamma^2 |\Phi_{\odot} \Delta d_{SE} [\text{AU}]|} .$$

Following [Altschul, 2009], we consider the beam energy measurements of the LEP Energy Working Group for the LEP 2 programme in the last few years of LEP operation. Analysis of their data on the synchrotron losses collected at different moments of time leads to the limit $|\Delta m_e/m_e| < 0.0013$, i.e., 0.1%, for both electrons and positrons.

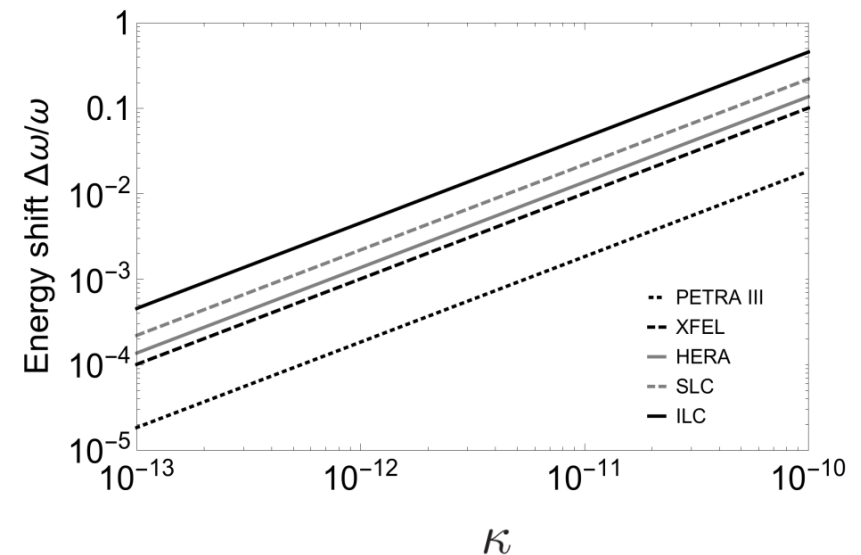
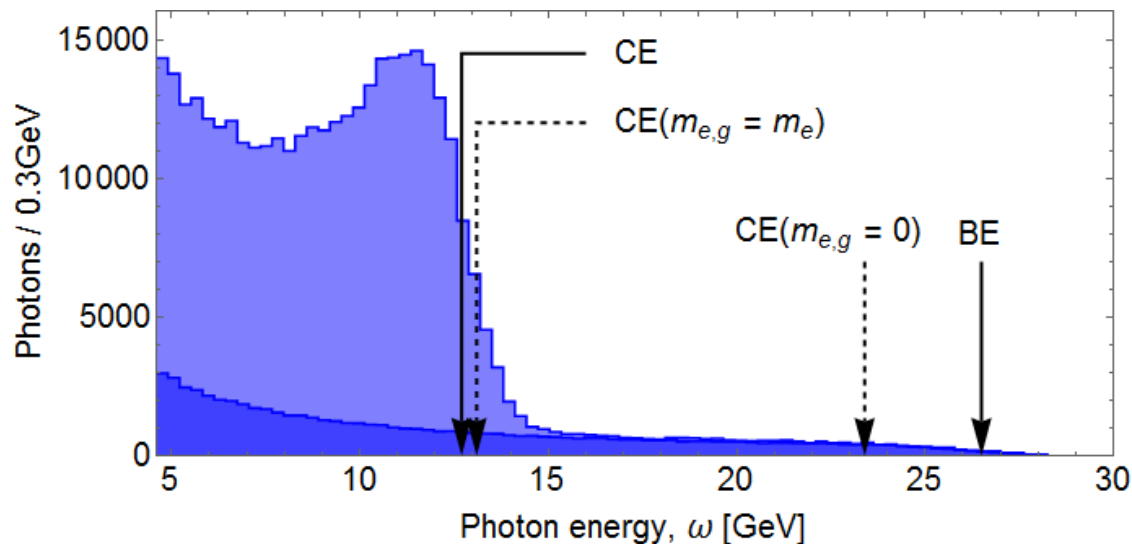
4. Compton scattering



Shifts in the Compton edge give

$$\left| \frac{\Delta m_e}{m_e} \right| < \frac{\Delta\omega_1 + \Delta\omega_2}{\omega_{max}} \cdot \frac{m^2(1+x)^2}{4\mathcal{E}^2|\Delta\Phi|}$$

where $x \equiv 4\mathcal{E}\omega_0 \sin^2(\theta_0/2)/m_e^2$



Results

- Absence of vacuum Cherenkov radiation at LEP and photon stability at Tevatron give 4% limit on the difference between the gravitational and inertial masses of the electron/positron at GeV energies.
- Synchrotron losses at LEP reduce this figure to 0.13%
- Compton scattering can provide a similar or better precision if performed at ILC/CLIC twice: when Earth is at the aphelion and perihelion of its orbit. One can also study day to day variations.

With development of accelerator technologies, we are finally able to rule out antigravity and confirm weak equivalence principle for the high-energy matter and antimatter.