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T.K., I. Kirsch, Phys.Rev.Lett. 106 (2011) 211601
I. Gahramanov, T.K., I. Kirsch, Phys.Rev. D85 (2012) 126013
T.K., "Chiral superfluidity of the quark-gluon plasma", ArXiv: 1208.0012
```

# Local parity violation in the quark-gluon plasma and its holographic description



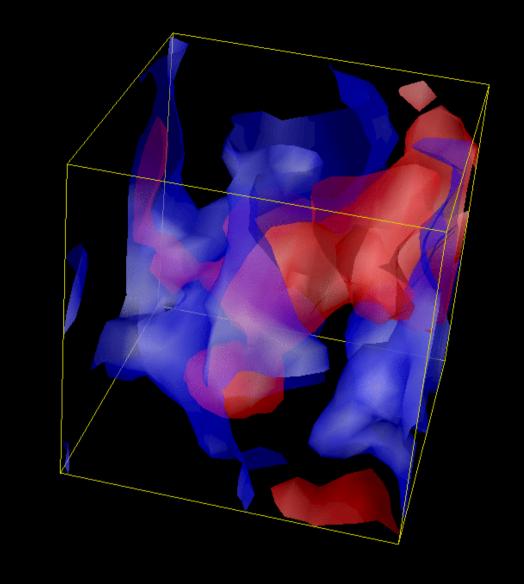
Tigran Kalaydzhyan

## QCD vacuum

$$G^{a\mu
u} ilde{G}^a_{\mu
u}$$



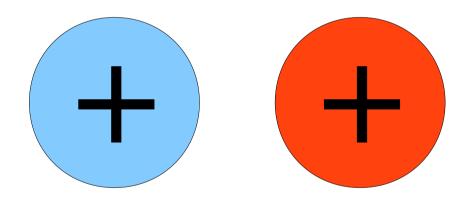
$$\rho_R \neq \rho_L$$

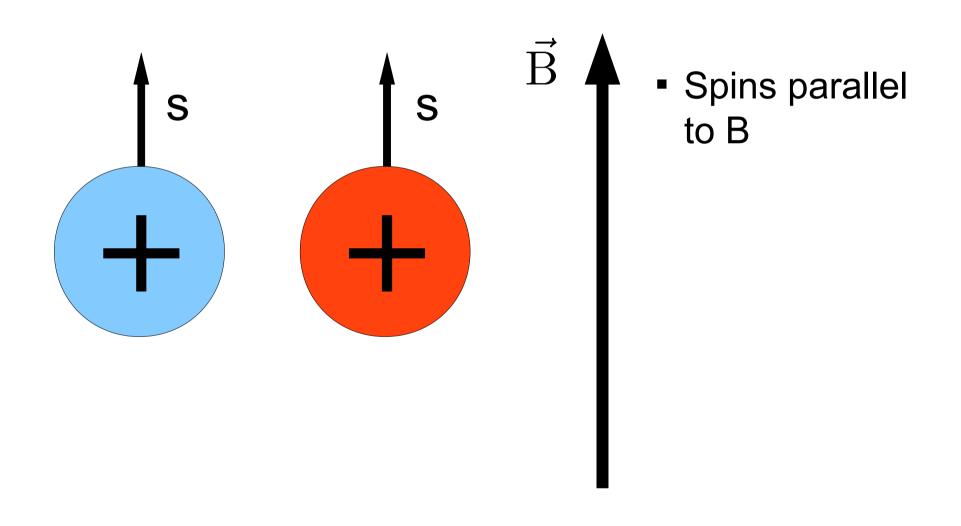


Positive topological charge density

Negative topological charge density

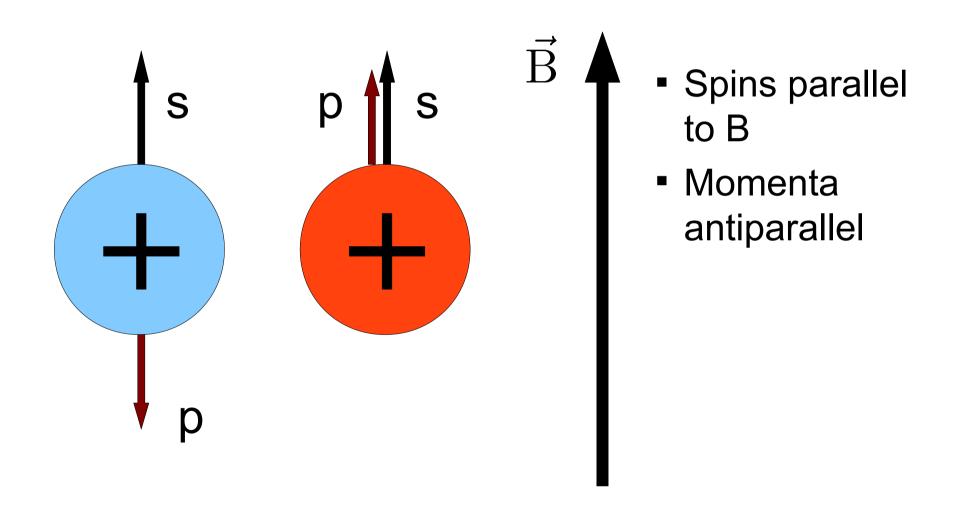
For the details of the simulation see P. Buividovich, T.K., M. Polikarpov PRD 86, 074511





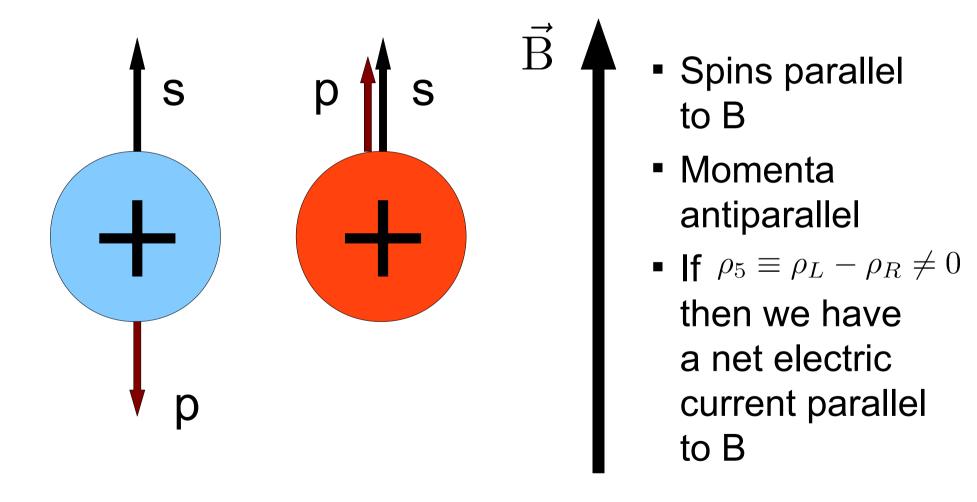
Right-handed

Left-handed



Right-handed

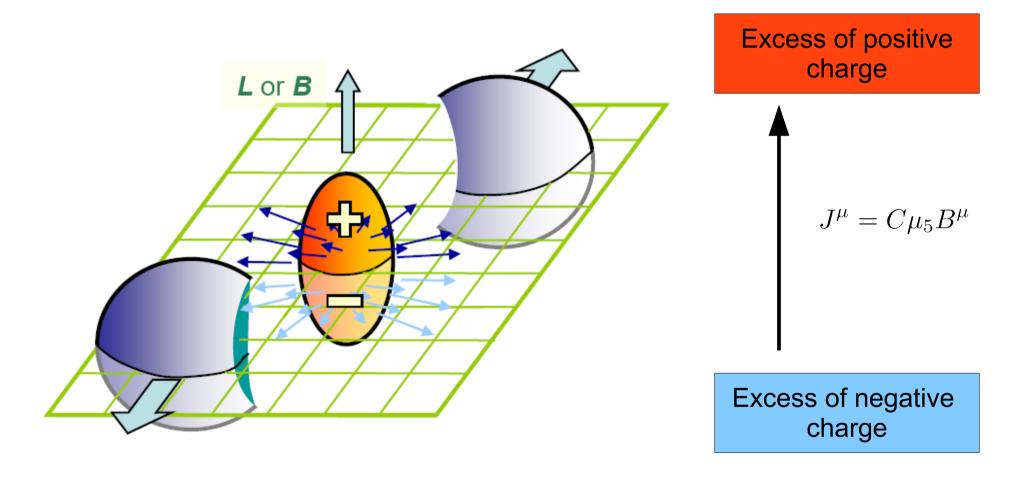
Left-handed



Left-handed

**Right-handed** 

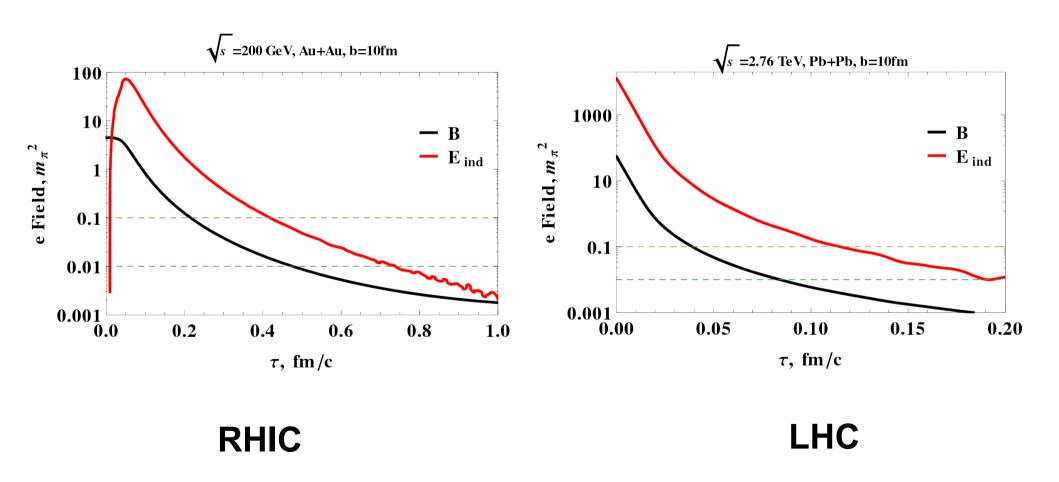
# Chiral Magnetic Effect



Fukushima, Kharzeev, McLerran, Warringa (2007)

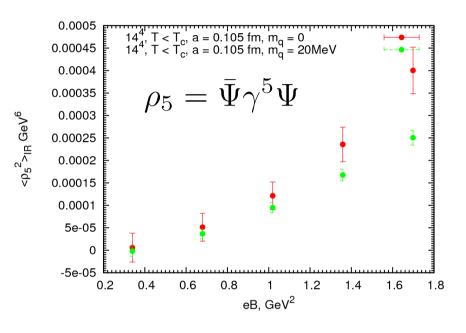
For a local strong parity violation see e.g. 0909.1717 (STAR) and 1207.0900 (ALICE)

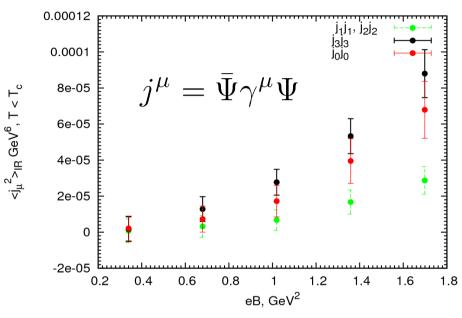
# Electromagnetic fields

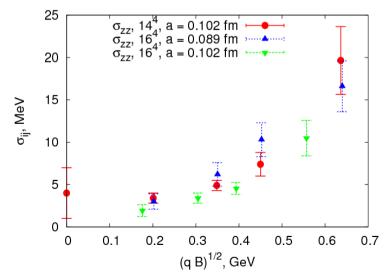


Huge electromagnetic fields, never observed before!

## Some numbers (lattice)



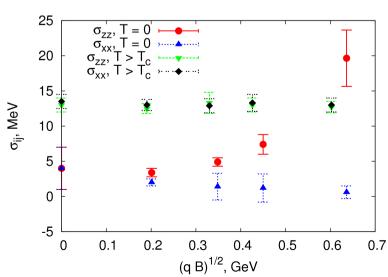




T.K., D. Kharzeev and

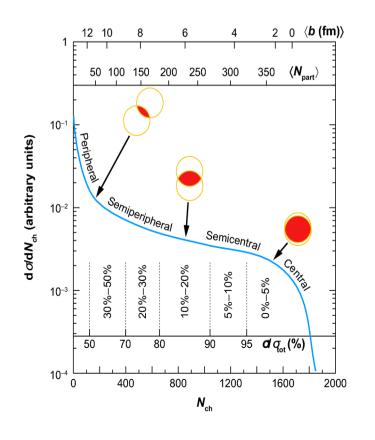


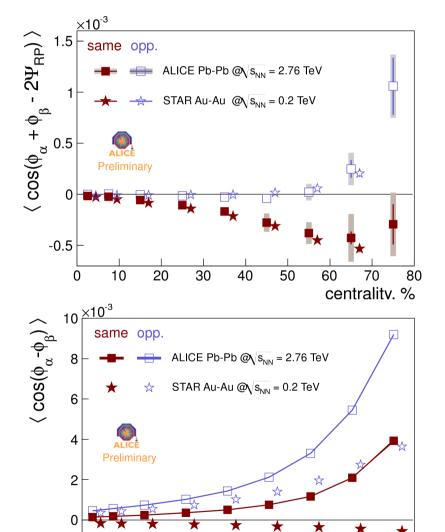
PRL 105 (2010) 132001 Phys.Atom.Nucl. 75, 488



$$\gamma_{\alpha,\beta} = \langle \cos(\phi_{\alpha} + \phi_{\beta}) \rangle = \langle \cos \cos \rangle - \langle \sin \sin \rangle$$

$$\delta_{\alpha,\beta} = \langle \cos(\phi_{\alpha} - \phi_{\beta}) \rangle = \langle \cos \cos \rangle + \langle \sin \sin \rangle$$

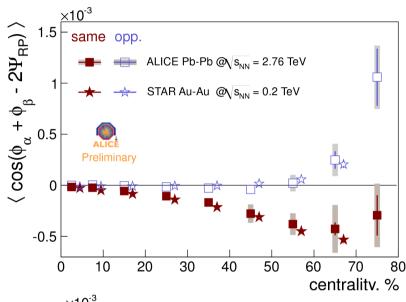


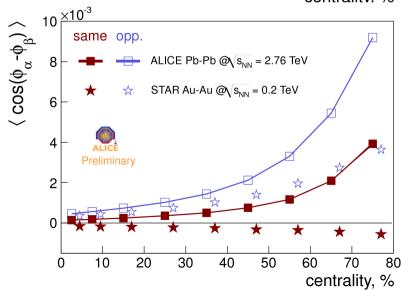


centrality, %

$$\gamma_{\alpha,\beta} = \langle \cos(\phi_{\alpha} + \phi_{\beta}) \rangle = \langle \cos\cos\rangle - \langle \sin\sin\rangle$$

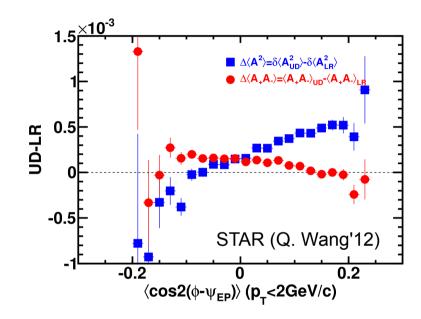
$$\delta_{\alpha,\beta} = \langle \cos(\phi_{\alpha} - \phi_{\beta}) \rangle = \langle \cos\cos\rangle + \langle \sin\sin\rangle$$
in-plane out-of-plane





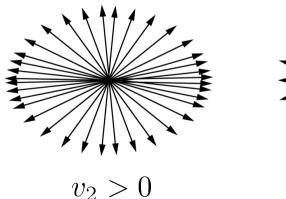
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in-plane out-of-plane



$$\gamma_{\alpha,\beta} \sim v_2 F_{\alpha,\beta} - H_{\alpha,\beta}^{\text{out}} + H_{\alpha,\beta}^{\text{in}}$$

$$\delta_{\alpha,\beta} \sim F_{\alpha,\beta} + H_{\alpha,\beta}^{\text{out}} + H_{\alpha,\beta}^{\text{in}}$$

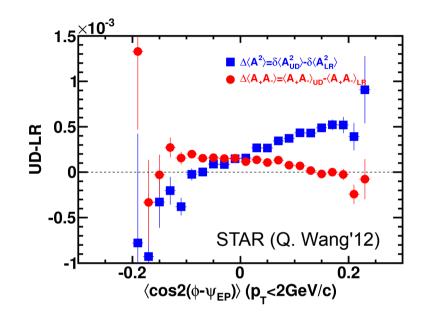


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in-plane out-of-plane

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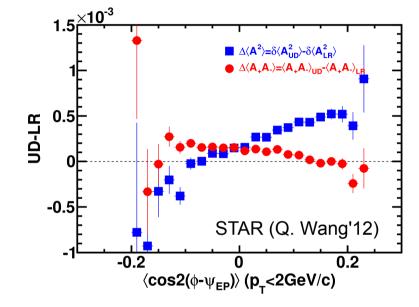


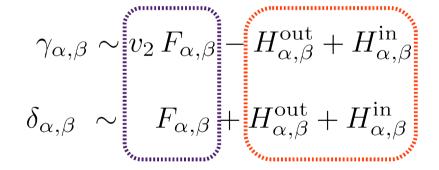
flow-dependent flow-independent

See also a nice review by Bzdak, Koch and Liao: ArXiv:1207.7327

$$\gamma_{\alpha,\beta} = \langle \cos(\phi_{\alpha} + \phi_{\beta}) \rangle = \langle \cos\cos\rangle - \langle \sin\sin\rangle$$

$$\delta_{\alpha,\beta} = \langle \cos(\phi_{\alpha} - \phi_{\beta}) \rangle = \langle \cos\cos\rangle + \langle \sin\sin\rangle$$
in-plane out-of-plane





flow-dependent flow-independent

From the data analysis it turns out that

$$|F_{\alpha,\beta}| \sim |H_{\alpha,\beta}|$$

#### **Questions**:

- Is CME flow dependent?
- What are other potential anomalous contributions?

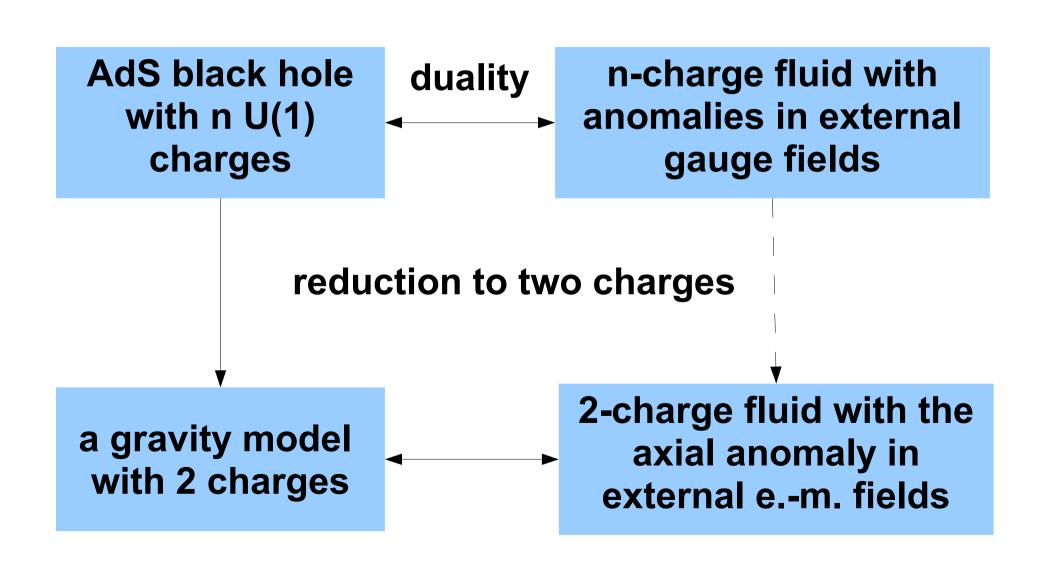
## Our task

- Find possible elliptic flow dependence of CME (in an optimistic assumption of a long-living magnetic field)
- Build a gravity dual to a strongly coupled relativistic anisotropic quantum fluid with the axial anomaly.
- Derive an effective model for QCD at Tc < T < 2 Tc (do we have anything in addition to CME?)
- Find hydrodynamic equations corresponding to the effective Lagrangian
- Extract phenomenological output for the heavy-ion collisions

# Elliptic flow dependence of

CME

## Main idea



## Hydrodynamics

#### Three-charge model:

$$\partial_{\mu}T^{\mu\nu}=F^{a\nu\lambda}j^{a}_{\lambda}$$
,

$$a = 1, 2, ..., n$$

$$\partial_{\mu} j^{a\mu} = -\frac{1}{8} C^{abc} F^{b}_{\mu\nu} \tilde{F}^{c\mu\nu} = C^{abc} E^{b} \cdot B^{c}_{\bullet}$$

#### **Electric field**

$$E^{a\mu} = u_{\nu} F^{a\mu\nu}$$

#### **Magnetic field**

$$B^{a\mu} = \frac{1}{2} \epsilon^{\mu\nu\lambda\rho} u_{\nu} F^{a}_{\lambda\rho}$$

# Hydrodynamics

#### Three-charge model:

$$\partial_{\mu}T^{\mu\nu}=F^{a\nu\lambda}j^{a}_{\lambda}$$
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#### where the stress-energy tensor and U(1) currents

$$T^{\mu\nu} = (\epsilon + P) u^{\mu} u^{\nu} + P g^{\mu\nu} + \dots,$$

$$j^{a\mu} = \rho^a u^{\mu} + \xi^a_{\omega} \omega^{\mu} + \xi^{ab}_{B} B^{b\mu} + \dots$$

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$$E^{a\mu} = u_{\nu} F^{a\mu\nu}$$

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$$B^{a\mu} = \frac{1}{2} \epsilon^{\mu\nu\lambda\rho} u_{\nu} F^{a}_{\lambda\rho}$$

#### **Vorticity**

$$\omega^{\mu} = \frac{1}{2} \epsilon^{\mu\nu\lambda\rho} u_{\nu} \partial_{\lambda} u_{\rho}$$

# Hydrodynamics

#### Three-charge model:

$$\partial_{\mu}T^{\mu\nu}=F^{a\nu\lambda}j^{a}_{\lambda}$$
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#### **Vorticity**

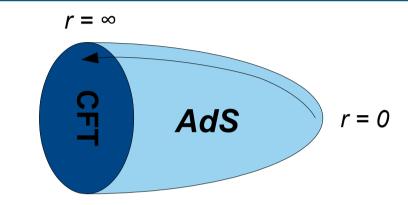
$$\omega^{\mu} = \frac{1}{2} \epsilon^{\mu\nu\lambda\rho} u_{\nu} \partial_{\lambda} u_{\rho}$$

Quantum anomaly → classical dynamics!

Son and Surowka (2009)

## Holography. Algorithm.

Fluid on the boundary, gravity in the bulk. Input: energy density, anomaly, background fields, etc.



Fix metric components (and gauge fields components), Chern-Simons parameters, etc in the bulk.

Solve equations of motion for the bulk fields (i.e. Einstein-Maxwell-Chern-Simons).

Read off a nontrivial result (i.e. transport coefficients) from the near-boundary expansion of the gravity solution.

see also Bhattacharyya, Hubeny, Minwalla and Rangamani (2008), Torabian and Yee (2009) Erdmenger, Haack, Kaminski, Yarom (2008)

#### Holograhic dual of conformal U(1)<sup>n</sup> theory:

$$\mathcal{L} = R - 2\Lambda - F_{MN}^a F^{aMN} + \frac{S_{abc}}{6\sqrt{-g}} \varepsilon^{PKLMN} A_P^a F_{KL}^b F_{MN}^c$$

Holograhic dual of conformal U(1)<sup>n</sup> theory:

logranic dual of conformal U(1) theory: 
$$S_{abc} = 4\pi G_5 C_{abc}$$
 
$$\mathcal{L} = R - 2\Lambda - F_{MN}^a F^{aMN} + \frac{S_{abc}}{6\sqrt{-g}} \varepsilon^{PKLMN} A_P^a F_{KL}^b F_{MN}^c$$

Holograhic dual of conformal U(1)<sup>n</sup> theory:

$$-S_{abc} = 4\pi G_5 C_{abc}$$

$$\mathcal{L} = R - 2\Lambda - F_{MN}^a F^{aMN} + \frac{S_{abc}}{6\sqrt{-g}} \varepsilon^{PKLMN} A_P^a F_{KL}^b F_{MN}^c$$

#### **Boosted AdS black hole solution:**

$$ds^{2} = -f(r)u_{\mu}u_{\nu}dx^{\mu}dx^{\nu} - 2u_{\mu}dx^{\mu}dr + r^{2}(\eta_{\mu\nu} + u_{\mu}u_{\nu})dx^{\mu}dx^{\nu}$$
$$A^{a} = (A_{0}^{a}(r)u_{\mu} + A_{\mu}^{a})dx^{\mu}$$

$$f(r) = r^2 - rac{m}{r^2} + \sum_a rac{(q^a)^2}{r^4}$$
 and  $A_0^a(r) = -rac{\sqrt{3}q^a}{2r^2}$ 

Holograhic dual of conformal U(1)<sup>n</sup> theory:

$$S_{abc} = 4\pi G_5 C_{abc}$$

**External electromagnetic fields** 

$$\mathcal{L} = R - 2\Lambda - F_{MN}^a F^{aMN} + \frac{S_{abc}}{6\sqrt{-g}} \varepsilon^{PKLMN} A_P^a F_{KL}^b F_{MN}^c$$

#### **Boosted AdS black hole solution:**

$$ds^2=-f(r)u_\mu u_\nu dx^\mu dx^\nu-2u_\mu dx^\mu dr+r^2(\eta_{\mu\nu}+u_\mu u_\nu)dx^\mu dx^\nu$$
 
$$A^a=(A_0^a(r)u_\mu+\mathcal{A}_\mu^a)dx^\mu$$
 4-velocity of the fluid

#### where

$$f(r)=r^2-rac{m}{r^2}+\sum_arac{(q^a)^2}{r^4}$$
 and  $A_0^a(r)=-rac{\sqrt{3}q^a}{2r^2}$  U(1) charges

Holograhic dual of conformal U(1)<sup>n</sup> theory:

$$-S_{abc} = 4\pi G_5 C_{abc}$$

$$\mathcal{L} = R - 2\Lambda - F_{MN}^a F^{aMN} + \frac{S_{abc}}{6\sqrt{-g}} \varepsilon^{PKLMN} A_P^a F_{KL}^b F_{MN}^c$$

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$$A^a = (A_0^a(r)u_\mu + A_\mu^a) dx^\mu$$
 External electromagnetic fields

where

$$\frac{f(r)=r^2-\frac{m}{r^2}+\sum_a\frac{(q^a)^2}{r^4}}{\qquad}\text{and}\qquad \frac{A_0^a(r)=-\frac{\sqrt{3}q^a}{2r^2}}{\qquad}\text{U(1) charges}$$

Hawking temperature:  $T \propto r_+$  Charge density:  $ho^a \propto q^a$ 

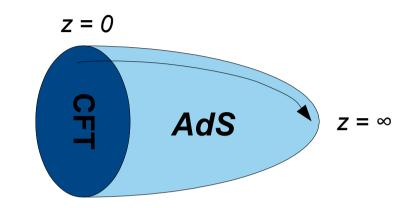
Chemical potentials:  $\mu^a \equiv A_0^a(r_+) - A_0^a(\infty)$  Pressure:  $P = \frac{\epsilon}{3} \propto m$ 

## Next order

We slowly vary 4-velocity and background fields

$$u_{\mu} = (-1, x^{\nu} \partial_{\nu} u_i)$$

$$\mathcal{A}^a_\mu = (0, x^\nu \partial_\nu \mathcal{A}^a_\mu)$$



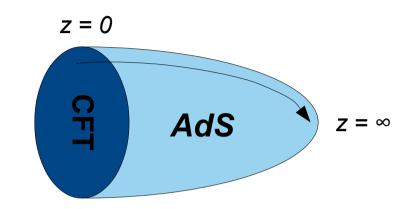
Then solve equations of motion for this case and find corrections to the metric and gauge fields.

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Then solve equations of motion for this case and find corrections to the metric and gauge fields.

And consider the near-boundary expansion (Fefferman-Graham coordinates):

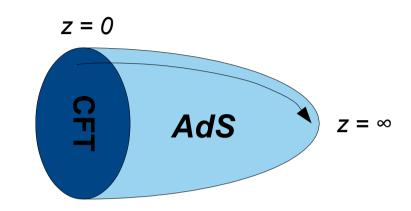
$$ds^{2} = \frac{1}{z^{2}} (g_{\mu\nu}(z, x) dx^{\mu} dx^{\nu} + dz^{2}),$$

## Next order

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$$u_{\mu} = (-1, x^{\nu} \partial_{\nu} u_i)$$

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$$ds^{2} = \frac{1}{z^{2}} (g_{\mu\nu}(z, x) dx^{\mu} dx^{\nu} + dz^{2}),$$

$$g_{\mu\nu}(z,x) = \eta_{\mu\nu} + g_{\mu\nu}^{(2)}(x)z^2 + g_{\mu\nu}^{(4)}(x)z^4 + \dots$$

$$A^a_\mu(z,x) = \mathcal{A}^a_\mu(x) + A^{a(2)}_\mu(x)z^2 + \dots$$

$$T_{\mu\nu} = \frac{g_{\mu\nu}^{(4)}(x)}{4\pi G_5} + \dots$$

$$j_a^{\mu} = \frac{\eta^{\mu\nu} A_{a\nu}^{(2)}(x)}{8\pi G_5} + \dots$$

## Transport coefficients

$$j^{a\mu} = \rho^a u^{\mu} + \xi^a_{\omega} \omega^{\mu} + \xi^{ab}_{B} B^{b\mu} + \dots$$

where the coefficients are

$$\xi_{\omega}^{a} = C^{abc} \mu^{b} \mu^{c} - \frac{2}{3} \rho^{a} C^{bcd} \frac{\mu^{b} \mu^{c} \mu^{d}}{\epsilon + P} + O(T^{2})$$

$$\xi_{B}^{ab} = C^{abc} \mu^{c} - \frac{1}{2} \rho^{a} C^{bcd} \frac{\mu^{c} \mu^{d}}{\epsilon + P} + O(T^{2})$$

Here  $\mu^a$  is a chemical potential associated with density  $\rho^a$ 

## Reduction to two charges

#### **Hydrodinamic equations:**

$$\partial_{\mu} T^{\mu\nu} = F^{\nu\lambda} j_{\lambda},$$

$$\partial_{\mu} j_{5}^{\mu} = C E^{\lambda} \cdot B_{\lambda} + \frac{C}{3} E_{5}^{\lambda} \cdot B_{5\lambda},$$

$$\partial_{\mu} j^{\mu} = 0$$

where vector and axial currents are

# $j^{\mu} = \rho u^{\mu} + \kappa_{\omega} \omega^{\mu} + \kappa_{B} B^{\mu} + \dots$ $j^{\mu}_{5} = \rho_{5} u^{\mu} + \xi_{\omega} \omega^{\mu} + \xi_{B} B^{\mu} + \dots$

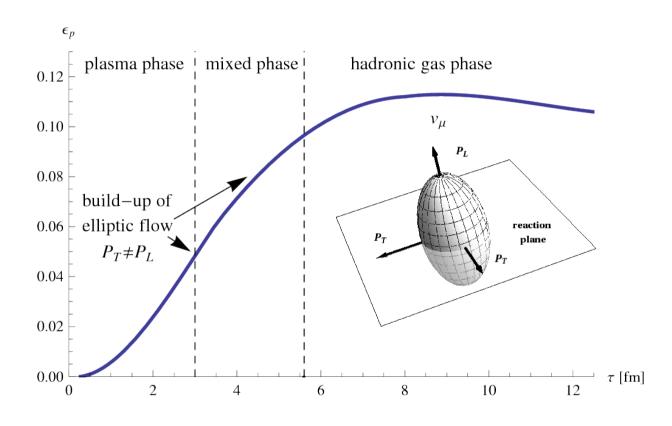
**Anomalies:** 

CVE 
$$\kappa_{\omega} = 2C\mu\mu_{5} \left(1 - \frac{\mu\rho}{\epsilon + P} \left[1 + \frac{\mu_{5}^{2}}{3\mu^{2}}\right]\right), \qquad \kappa_{B} = C\mu_{5} \left(1 - \frac{\mu\rho}{\epsilon + P}\right), \qquad \text{CME}$$

$$\xi_{\omega} = C\mu^{2} \left(1 - 2\frac{\mu_{5}\rho_{5}}{\epsilon + P} \left[1 + \frac{\mu_{5}^{2}}{3\mu^{2}}\right]\right), \qquad \xi_{B} = C\mu \left(1 - \frac{\mu_{5}\rho_{5}}{\epsilon + P}\right), \qquad \text{CSE}$$

T.K. and I. Kirsch, PRL 106 (2011) 211601 + PRD 85 (2012) 126013

## Anisotropic case

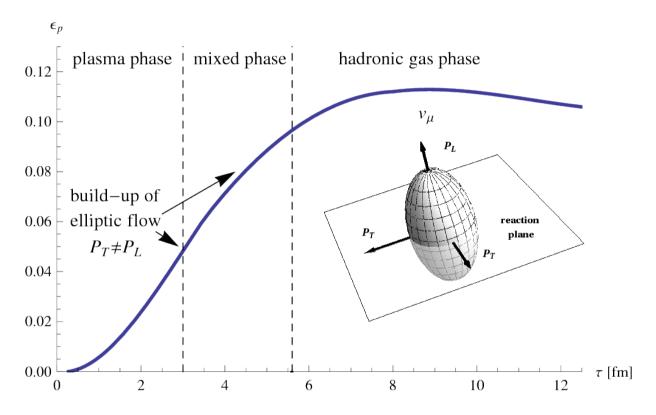


$$T^{\mu\nu} = \begin{pmatrix} \epsilon & 0 & 0 & 0 \\ 0 & P_T & 0 & 0 \\ 0 & 0 & P_T & 0 \\ 0 & 0 & 0 & P_L \end{pmatrix}$$

#### anisotropy parameter

$$\epsilon_P = \frac{P_T - P_L}{P_T + P_L}$$

## Anisotropic case



$$T^{\mu\nu} = \begin{pmatrix} \epsilon & 0 & 0 & 0 \\ 0 & P_T & 0 & 0 \\ 0 & 0 & P_T & 0 \\ 0 & 0 & 0 & P_L \end{pmatrix}$$

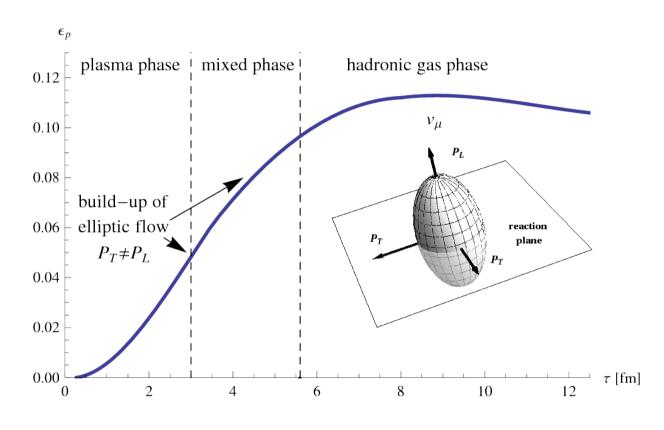
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#### can be translated at freeze-out to

$$v_2 \approx \epsilon_P/2$$

## Anisotropic case



$$T^{\mu
u} = \left( egin{array}{cccc} \epsilon & 0 & 0 & 0 \ 0 & P_T & 0 & 0 \ 0 & 0 & P_T & 0 \ 0 & 0 & 0 & P_L \end{array} 
ight)$$

anisotropy parameter

$$\epsilon_P = \frac{P_T - P_L}{P_T + P_L}$$

can be translated at freeze-out to

$$v_2 \approx \epsilon_P/2$$

$$T^{\mu\nu} = (\epsilon + P_T)u^{\mu}u^{\nu} + P_T g^{\mu\nu} - \Delta v^{\mu}v^{\nu} + \tau^{\mu\nu}$$

$$j^{a\mu} = \rho^a u^\mu + \nu^{a\mu}$$

$$j^{a\mu} = 
ho^a u^\mu + 
u^{a\mu}$$
 where  $u_\mu u^\mu = -1 \,, v_\mu v^\mu = 1 \,, u_\mu v^\mu = 0 \,.$ 

## Gravity side

#### Anisotropic AdS geometry with multiple U(1) charges:

$$ds^{2} = -f(r)u_{\mu}u_{\nu}dx^{\mu}dx^{\nu} - 2u_{\mu}dx^{\mu}dr$$

$$+r^{2}w_{T}(r)P_{\mu\nu}dx^{\mu}dx^{\nu} - r^{2}(w_{T}(r) - w_{L}(r))v_{\mu}v_{\nu}dx^{\mu}dx^{\nu}$$

$$A^{a} = (A_{0}^{a}(r)u_{\mu} + A_{\mu}^{a})dx^{\mu}$$

#### Where, close to the boundary,

$$f(r) = r^2 - \frac{m}{r^2} + \sum_{a} \frac{(q^a)^2}{r^4} + \mathcal{O}(r^{-6})$$
$$A_0^a(r) = \mu_\infty^a - \frac{\sqrt{3}q^a}{2r^2} + \mathcal{O}(r^{-8})$$

## Gravity side

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$$w_{T}(r) = 1 + \frac{m\zeta}{4r^{4}} + \mathcal{O}(r^{-8})$$

$$A_{0}^{a}(r) = \mu_{\infty}^{a} - \frac{\sqrt{3}q^{a}}{2r^{2}} + \mathcal{O}(r^{-8})$$

$$w_{L}(r) = 1 - \frac{m\zeta}{2r^{4}} + \mathcal{O}(r^{-8})$$

Parameter zeta is related to the anisotropy:

$$\zeta = \frac{2\epsilon_P}{\epsilon_P + 3}$$

### Anisotropic CME

In general one has to solve the EOM not only close to the boundary, but also deeper in the bulk, up to the horizon. By doing this (numerically) and reading off the transport coefficients, we get (to linear order in anisotropy)

$$\kappa_B = C\mu_5 \left( 1 - \frac{\mu\rho}{\epsilon + \bar{P}} \left[ 1 - \frac{\varepsilon_p}{6} \right] \right)$$

Where the average pressure

$$\bar{P} = \frac{2P_T + P_L}{3}$$

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$$\frac{\mu\rho}{\epsilon + \bar{P}} \approx \frac{\mu}{T} \lesssim \frac{10 \,\mathrm{MeV}}{200 \,\mathrm{MeV}} \ll 1$$

### Anisotropic CME

In general one has to solve the EOM not only close to the boundary, but also deeper in the bulk, up to the horizon. By doing this (numerically) and reading off the transport coefficients, we get (to linear order in anisotropy)

$$\kappa_B = C\mu_5 \left( 1 - \frac{\mu\rho}{\epsilon + \bar{P}} \left[ 1 - \frac{\varepsilon_p}{6} \right] \right)$$

Where the average pressure

$$\bar{P} = \frac{2P_T + P_L}{3}$$

$$\frac{\mu\rho}{\epsilon + \bar{P}} \approx \frac{\mu}{T} \lesssim \frac{10 \,\text{MeV}}{200 \,\text{MeV}} \ll 1$$

Taken that  $v_2 pprox \epsilon_P/2$  close to the hadronization we conclude, that

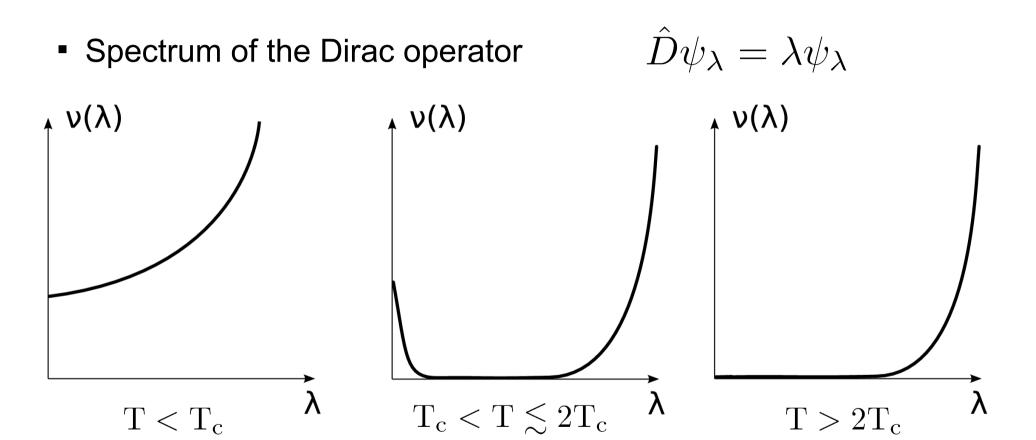
Chiral Magnetic Effect depends weakly on the elliptic flow and can be separated from the purely hydrodynamic effects!

# Parity-odd

### effects from the

first principles

### Insight from the lattice

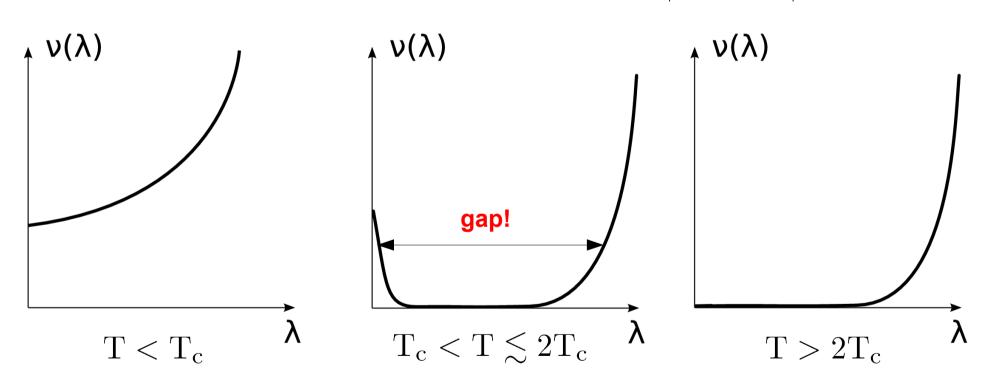


Chiral properties are described by near-zero modes

### Insight from the lattice

Spectrum of the Dirac operator

$$\hat{D}\psi_{\lambda} = \lambda\psi_{\lambda}$$



- Chiral properties are described by near-zero modes
- There are two separated parts of the spectrum at intermediate temperatures!

## Why "superfluidity"?

Energy Normal motions **Curl-free motions** 

AUGUST 15, 1941

PHYSICAL REVIEW

#### Theory of the Superfluidity of Helium II

L. LANDAU
Institute of Physical Problems, Academy of Sciences USSR, Moscow, USSR

Therefore, between the lowest energy levels of vortex and potential motion there must be a certain energy interval  $\Delta$ .

The supposition that the normal level of potential motions lies lower than the beginning of the spectrum of vortex motions leads to the phenomenon of superfluidity.

One of these motions is "normal" and the other is "superfluid."

We will not consider any spontaneously broken symmetry!

• Euclidean functional integral for  $SU(N_c) \times U_{em}(1)$  is given by

$$\int D\bar{\psi}D\psi \exp\left\{-\int_{V} d^{4}x \,\bar{\psi}(D\!\!\!\!/ - im)\psi + \frac{1}{4}G^{a\mu\nu}G^{a}_{\mu\nu} + \frac{1}{4}F^{\mu\nu}F_{\mu\nu}\right\},\,$$

$$D = -i(\partial + A + g G + \gamma_5 A_5).$$

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where we define the Dirac operator as

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- consider a pure gauge  $A_{5\mu}=\partial_{\mu}\theta$  for the auxiliary axial field
- and the chiral limit  $m \to 0$

The total effective Euclidean Lagrangian reads as

$$\mathcal{L}_{E}^{(4)} = \frac{1}{4} G^{a\mu\nu} G^{a}_{\mu\nu} + \frac{1}{4} F^{\mu\nu} F_{\mu\nu} - j^{\mu} A_{\mu}$$

$$+ \frac{\Lambda^{2} N_{c}}{4\pi^{2}} \partial^{\mu} \theta \partial_{\mu} \theta + \frac{g^{2}}{16\pi^{2}} \theta G^{a\mu\nu} \widetilde{G}^{a}_{\mu\nu} + \frac{N_{c}}{8\pi^{2}} \theta F^{\mu\nu} \widetilde{F}_{\mu\nu}$$

$$+ \frac{N_{c}}{24\pi^{2}} \theta \Box^{2} \theta - \frac{N_{c}}{12\pi^{2}} \left( \partial^{\mu} \theta \partial_{\mu} \theta \right)^{2}$$

So we get an axion-like field with decay constant  $f=\frac{2\Lambda}{\pi}\sqrt{N_c}$  and a negligible mass  $m_{\theta}^2=\lim_{V\to\infty}\frac{\langle Q^2\rangle}{f^2V}\equiv\chi(T)/f^2$  .

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Dynamical axion-like internal degree of freedom in QCD!

#### Interpretation of the scale \(\Lambda\)

• From the quartic Lagrangian at  $N_c=N_f=1$  we get

$$\rho_5 = -\lim_{t_E \to 0} \frac{\delta \mathcal{L}_E^{(4)}}{\delta \mu_5} = \frac{1}{2} \left(\frac{\Lambda}{\pi}\right)^2 \mu_5 + \frac{1}{3\pi^2} \mu_5^3$$

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- Free quarks and a strong B-field:  $\Lambda = 2\sqrt{|eB|}$
- Dynamical fermions (1105.0385):  $\Lambda \simeq 3\,\mathrm{GeV} \gg \Lambda_{QCD}$

Considering EOM for the Minkowski effective Lagrangian and only the color-singlet states, we obtain:

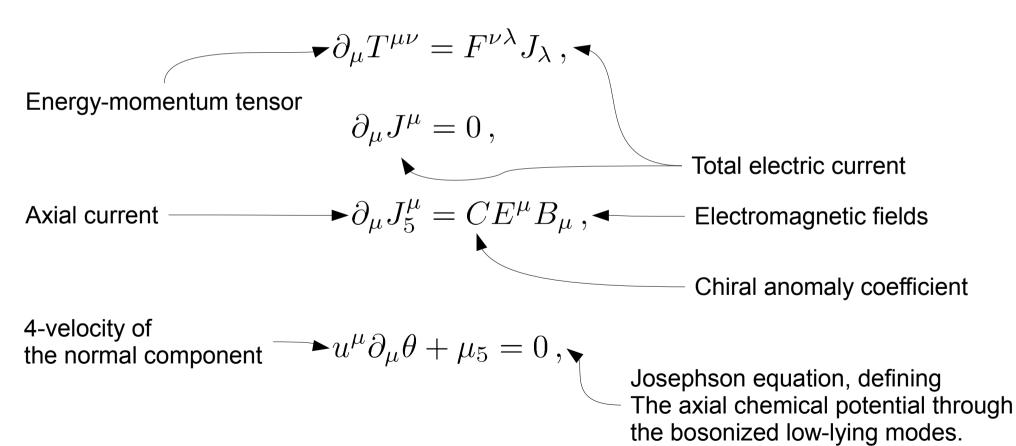
$$\partial_{\mu}T^{\mu\nu} = F^{\nu\lambda}J_{\lambda} \,,$$

$$\partial_{\mu}J^{\mu}=0\,,$$

$$\partial_{\mu}J_{5}^{\mu} = CE^{\mu}B_{\mu} \,,$$

$$u^{\mu}\partial_{\mu}\theta + \mu_5 = 0,$$

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Similar to the superfluid dynamics!

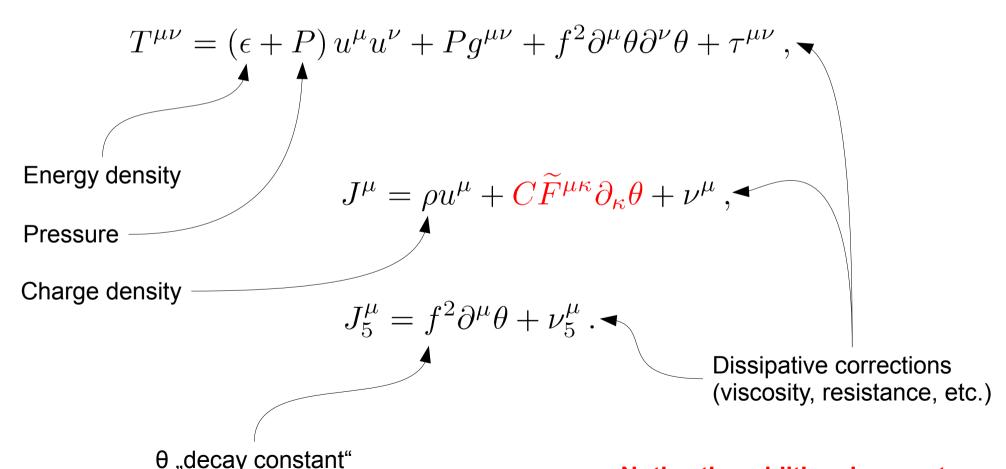
 Solving the hydrodynamic equations in the gradient expansion, we obtain the constitutive relations:

$$T^{\mu\nu} = (\epsilon + P) u^{\mu}u^{\nu} + Pg^{\mu\nu} + f^2 \partial^{\mu}\theta \partial^{\nu}\theta + \tau^{\mu\nu} ,$$

$$J^{\mu} = \rho u^{\mu} + C\widetilde{F}^{\mu\kappa} \partial_{\kappa} \theta + \nu^{\mu} ,$$

$$J_5^{\mu} = f^2 \partial^{\mu} \theta + \nu_5^{\mu} .$$

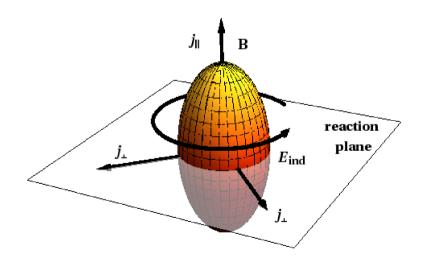
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Notice the additional current

An additional electric current induced by the  $\theta$ -field:

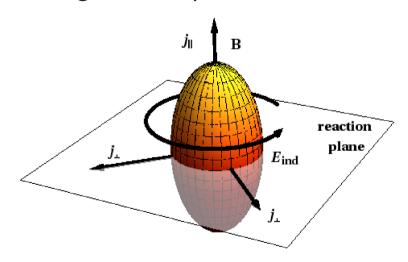
$$j_{\lambda} = -C\mu_5 B_{\lambda} + C\epsilon_{\lambda\alpha\kappa\beta} u^{\alpha} \partial^{\kappa} \theta E^{\beta} - Cu_{\lambda} (\partial \theta \cdot B)$$



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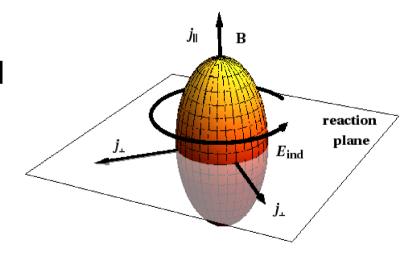
Chiral Magnetic Effect (electric current along B-field)



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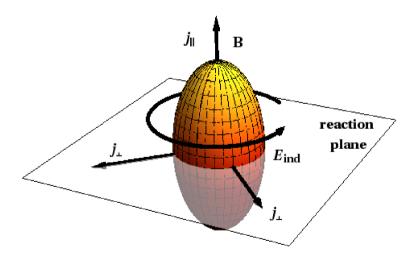
- Chiral Magnetic Effect (electric current along B-field)
- Chiral Electric Effect (electric current transverse to E-field and to both normal and superfluid velocities)



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- Chiral Dipole Wave (dipole moment induced by B-field)



An additional electric current induced by the  $\theta$ -field:

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reaction plane

- Chiral Magnetic Effect (electric current along B-field)
- Chiral Electric Effect (electric current transverse to E-field and to both normal and superfluid velocities)
- Chiral Dipole Wave (dipole moment induced by B-field)
- The field θ(x) itself: Chiral Magnetic Wave
   (propagating imbalance between the number of left- and right-handed quarks)

### Interesting projects

- Add more flavors. The "axion-like" field might be just a pion. At strong B it will be then a 2D pion.
- Role of the low-dimensional defects in inducing superfluidity/superconductivity. Copenhagen vacuum.
- Entropy and thermalization of various parts of fermionic spectrum. Relation to the vacuum entropy.
- Considering vortices (axionic strings). One might reproduce the chiral vortical effect.
- Holographic model of the chiral superfluidity and highorder corrections.
- Experimental searches for the mentioned effects.
- Chiral electric effect on a lattice.

#### Thank you for the attention!

and

### Have a good time!

All comments on the papers are welcome!
Also feel free to ask questions about the experimental observables.

#### One more remark

"Axionic" part of the Lagrangian

$$\mathcal{L}_{\theta} = \frac{\Lambda^2 N_c}{4\pi^2} \partial^{\mu}\theta \partial_{\mu}\theta + \frac{N_c}{24\pi^2} \theta \Box^2 \theta - \frac{N_c}{12\pi^2} \left(\partial^{\mu}\theta \partial_{\mu}\theta\right)^2 + \dots$$

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If magnetic field dominates over other scales, then we can make the following redefinition:  $\theta \to \frac{\pi}{\sqrt{2N_c eB}}\theta$ 

$$\mathcal{L}_{\theta} \to \frac{1}{2} \partial^{\mu} \theta \partial_{\mu} \theta + \frac{1}{48 eB} \theta \Box^{2} \theta - \frac{\pi^{2}}{48 N_{c}(eB)^{2}} (\partial^{\mu} \theta \partial_{\mu} \theta)^{2} + \dots$$

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In the limit  $B \to \infty$  bosonization becomes exact, which is an evidence of the  $(3+1) \to (1+1)$  reduction!